The 'Geometric Weak Gravity Conjecture'

(plus: The WGC and Axion Monodromy)

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based on work with F. Rompineve and A. Westphal

<u>Outline</u>

- Personal motivation
- The WGC without (useful) extremal objects
- The use of dualitites
- The geometric WGC

if time permits:

• Constraining monodromy by the WGC for domain walls

 The WGC is interesting as ...

Arkani-Hamed/Motl/Nicolis/Vafa '06

1) A possible fundamental feature of quantum gravity

• It quantifies the non-existence of global symmetries

(If $g \rightarrow 0$ is impossible, we need to know g_{min} . The WGC states $g_{min} = m$.)

- It may define a non-trivial boundary between landscape and swampland.
- Since it's always respected by string theory, it may teach us about the string's interplay with 'generic' quantum gravity.
- It may relate very directly to phenomenology....

The WGC is interesting because ...

2) The inflationary tensor-to-scalar ratio is...

$$r \equiv \frac{\Delta_T^2}{\Delta_R^2} = 16\epsilon \simeq 8 \left| \frac{d\varphi}{dN} \right|^2 \quad \Rightarrow \quad \Delta \varphi \simeq 20\sqrt{r} \,,$$

assuming $N \simeq 60$. (This is known as the Lyth bound).

- Thus, even though the BICEP 'discovery' of $r \simeq 0.15$ went away, the need to consider large-field models may return.
- Note: The Planck/BICEP analysis still sees a ($\sim 1.8\sigma$) hint for $r \simeq 0.05$. Much better values/bounds are expected soon.

Cheung/Remmen; de la Fuente/Saraswat/Sundrum ...'14 Rudelius; Ibanez/Montero/Uranga/Valenzuela; Brown/Cottrell/Shiu/Soler; Bachlechner/Long/McAllister; AH/Mangat/Rompineve/Witkowski; Junghans; Heidenreich/Reece/Rudelius; Kooner/Parameswaran/Zavala; Harlow; AH/Rompineve/Westphal; ...'15; Conlon/Krippendorf ...'16 The WGC is interesting because ...

3) It has the potential to constrain Relaxion models...

Ibanez/Montero/Uranga/Valenzuela '15

• Such models fine-tune the Higgs mass-squared dynamically (during inflation), but require a large field range of an axionic scalar to do so...

Graham/Kaplan/Rajendran '15



Crucially, one needs to constrain monodromy models Brown/Cottrell/Shiu/Soler, Ibanez/Montero/Uranga/Valenzuela '15 cf. also this and A. Westphal's talk

Fig. from Jaeckel/Metha/Witkowski '15

The (generalized) weak gravity conjecture

• The basic underlying lagrangian is (for *p*-dim. objects in *d* dims.; with $\overline{M}_P \equiv 1$)

$$S \sim \frac{1}{g^2} \int (F_{p+1})^2 + T \int_{p-dim.} dV + \int_{p-dim.} A_p$$

with

$$F_{p+1}=dA_p$$
 .

• To avoid stable extremal black branes, one requires charged objects with sub-extremal mass (tension):

$$q/T \geq \gamma_{p,d}^{1/2}$$
, where $\gamma_{p,d} = rac{p(d-p-2)}{d-2}$.

• As one clearly sees, this fails for instantons and objects with codimension 1 & 2 (domain walls and cosmic 'strings').

Heidenreich/Reece/Rudelius '15

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Note:

- This failure outside the range 0 is not unexpected:
- Indeed, the argument that 'the WGC protects us from too many stable objects' fails also intuitively outside this range.
 (E.g. strings and domain walls induce see e.g. Susskind '95

(E.g., strings and domain walls induce no long-range gravitational force.)

However:

 The arguments that 'the WGC protects us from the global-symmetry limit' and the investment of the MGCC'

'string theory always obeys the WGC' support the conjecture even outside the above range.

- Arguments supporting/quantifying the WGC outside the 'canonical range' of 0
 - string dualities

- consistency of generic KK-reductions
- consideration of dilatonic black branes.

Heidenreich/Reece/Rudelius '15

• Example for duality argument:

IIB on $\mathsf{CY}_3\times \mathbb{R}^3\times {{S}^1}$ — brane wrapped to give instantons

 \Leftarrow T-duality \Rightarrow

IIA on $CY_3 \times \mathbb{R}^3 \times S^{1'}$ — brane wrapped to give particle (at large R' and large coupling; hence M-theory in 5d; WGC applicable to 5d Reissner-Nordstroem-BHs).

 \Rightarrow WGC is carried back to IIB axion decay constant.

Brown/Cottrell/Shiu/Soler '15

- In fact, the key is not in the dualities, but rather in the same CY underlying both the M-theory and the IIB model.
- Hence, there ought to be a

Geometric WGC

- Consider a IIA-CY X with D2-branes wrapped on 2-cycles.
- Let w_i be a basis of $H^2(X, \mathbb{Z})$.

The metric on X induces a metric for 2-forms,

$$\mathcal{K}_{ij}\equiv\int_X w_i\wedge\star w_j\,,$$

and on the (dual) space of 2-cycles, K^{ij} .

• We make the standard ansatz

$$C_3 = A_1^i(x) \wedge w_i(y).$$

- Focus on 4d particles coming from D2s on a particular cycle Σ
- The relevant 4d action reads $(I_s = 1)$

$$S_4 \sim (V_X/g_s^2) \int \sqrt{g} R + \int K_{ij} F_2^i \wedge \star F_2^j + q_i^{\Sigma} \int A_1^i$$
world-line

with the charges

$$q_i^{\Sigma} = \int_{\Sigma} w^i$$
.

 Note that only a particular combination of Aⁱ₁'s is sourced by particles 'from Σ':

$$A_1^i \equiv A_1 \, K^{ij} q_i^{\Sigma}$$
 (this defines A_1).

• Thus, one arrives at the standard action

$$S_4 \supset rac{1}{2e^2} \int F_2 \wedge \star F_2 + \int A_1 \, ,$$
 world-line

with e^2 given by.....

• With e^2 given by

$$e^2 = 2\pi |q^{\Sigma}|^2$$
 with $|q^{\Sigma}|^2 \equiv K^{ij} q_i^{\Sigma} q_i^{\Sigma}$.

Here we reinstanted $\mathcal{O}(1)$ factors.

<u>Note:</u> Metric on $X \rightarrow$ natural norm on *p*-form space \rightarrow natural norm $|q^{\Sigma}|$ on *p*-cycle space.

• Finally, use $\overline{M}_P^2 = V_X / \kappa_{10}^2 g_s^2$ together with $M_{\Sigma} = (\mu_2/g_s) V_{\Sigma}$ and impose the WGC:

$$rac{e\overline{M}_P}{M_\Sigma} \geq rac{1}{\sqrt{2}} \qquad \Rightarrow \qquad rac{|q^\Sigma| \, V_X^{1/2}}{V_\Sigma} \geq rac{1}{2} \, .$$

Thus a particular, purely geometric (rescaling- and g_s -independent) quantity characterizing X is constrained.

- Crucially, the same function appears in WGC constraints on other objects obtained from other branes wrapped on 2-cycles.
- For example, D4s give domain walls with

$$\frac{e_{DW}\overline{M}_P}{T_{DW}} = \frac{(2V_X)^{1/2}|q^{\Sigma}|}{V_{\Sigma}}.$$

• Thus, using the 'particle-WGC', we constrain $V_X^{1/2}|q^{\Sigma}|/V_{\Sigma}$, obtaining a precise 'domain-wall-WGC':

$$\frac{e_{DW}\overline{M}_P}{T_{DW}} \geq \frac{1}{2} \,.$$

• This goes through for any dimension of the cycle Σ and any dimension of the brane. Hence, any object in 4d is constrained by the imposition of the WGC for particles.

• Thus, allowing also for multiple gauge fields,

Cheung/Remmen '14; Rudelius '14/'15, Brown/Cottrell/Shiu/Soler, Bachlechner/Long/McAllister '15

we find in full generality:

Geometric conjecture:

The convex hull spanned by the vectors $(V_X^{1/2}/V_{\Sigma}) q^{\Sigma}$

(with $\Sigma \in H^p(X, \mathbb{Z})$) contains the ball of radius 1/2.

Implication for (q+1)-dimensional objects in 4d:

The convex hull spanned by the vectors $(\overline{M}_P/T_q) \tilde{q}^{\Sigma}$ contains the ball of radius $1/\sqrt{2}$.

• Note: We did not use SUSY, the CY-condition, or the existence of a SUSY-brane on Σ . So this may be much stronger then the 'not too surprising' BPS-like result.

Constraining axion monodromy with the WGC

Disclaimer:

Only brief summary; for deeper analysis and relation to earlier work...

Kaloper/Lawrence/Sorbo '08..'11 (see also Dvali '05) Brown/Cottrell/Shiu/Soler; Ibanez/Montero/Uranga/Valenzuela '15 see talks by I. Valenzuela and A. Westphal.

- Let's assume, based on the above, that all 4d objects, in particular DWs, are constrained.
- <u>Note</u>: the 'light' stringy objects fulfilling the WGC above are nevertheless always heavier than the KK-scale $M_{KK} = \Lambda$.
- Thus, one might conjecture that the magnetic WGC

 $\Lambda^3 \lesssim e_2 \overline{M}_P$

always holds.

 Start from the 'standard' monodromy potential (with 'instantonic wiggles')
 AH/Rompineve/Westphal '15



(Effective) domain walls are automatically present, but are too light to give any useful WGC constraint.

(In fact, one may argue that they make the electric WGC useless.)

• Nevertheless, the effective action

$$S \sim \int \frac{1}{2(e_2)^2} F_4^2 + \int_{DW} A_3$$

is there and, using the quantization $F_4 = n e_2^2$, allows for matching the discrete effective potential

$$V(F_4)_{eff} = \frac{1}{2}(e_2)^2 n^2$$

to the previous effective potential

$$V(\varphi)_{eff}=\frac{1}{2}m^2(2\pi nf)^2\,.$$

• This implies $e_2 = 2\pi m f$ and hence

$$\Lambda^3 \lesssim e_2 \overline{M}_P = 2\pi m f \overline{M}_P$$
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• In the context of inflation, one has

$$H \sim m \varphi_{max} \lesssim \Lambda$$

and hence

$$\Lambda^3 \sim m f \, \overline{M}_P \qquad \Rightarrow \qquad \frac{\varphi_{max}}{\overline{M}_P} \lesssim \left(\frac{\overline{M}_P}{m}\right)^{2/3} \left(\frac{2\pi f}{\overline{M}_P}\right)^{1/3} \, .$$

To be better explained in A. Westphal's talk....

Summary

- Let's assume that string compactifications with form-fields / wrapped objects always obey the particle WGC.
- Then a geometric WGC follows.
- From this, one obtains a generalized WGC including axions, cosmic strings and DWs etc.
- The KK scale is always so low that also the generalized magnetic WGC is holds.
 Let's accept this latter form also more generally.
- The magnetic WGC for DWs provides for a very direct way of constraining axion-monodromy-type scalar potentials.