### Shift symmetries for the Higgs and the inflaton

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# Outline

• The 125-GeV-Higgs (without SUSY) from a String-Pheno Perspective

 $\rightarrow$  1204.2551 with A. Knochel and T. Weigand

- <u>Main idea</u>:  $\lambda = 0$  at some high scale (SUSY-breaking scale) due to shift symmetry in the Higgs sector
- Stringy origin of this shift symmetry
- <u>Closely related</u>: The very same symmetry may be reponsible for a flat potential in fluxbrane inflation

## Motivation

• We have a Higgs at 125 GeV and nothing else (yet?)

Of course: low-scale SUSY is still OK Also: Muon-(g - 2);  $h \rightarrow \gamma \gamma$  excess; 130-GeV  $\gamma$ -ray line...

- <u>Nevertheless</u>: What if we just had to accept the fine-tuned non-SUSY SM for a large energy range?
- <u>Well-known</u>: for low  $m_h$ ,  $\lambda$  runs to zero at some scale  $< M_P$  (vacuum stability bound)

. . .

Lindner, Sher, Zaglauer '89 Gogoladze, Okada, Shafi '07

Shaposhnikov, Wetterich 09' Giudice, Isidori, Strumia, Riotto, ...

• It has been attempted to turn this into an  $m_h$  prediction



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### String-phenomenologist's perspective

- Insist on stringy UV completion (for conceptual reasons)
- Expect SUSY at string/compactification scale (stability!)
- Natural guess: The special scale  $\mu(\lambda = 0)$  is the SUSY-breaking scale
- Crucial formula:

$$\lambda(m_s) = \frac{g^2(m_s) + g'^2(m_s)}{8} \cos^2(2\beta)$$

• Reminder:

$$M_{H}^{2} = \begin{pmatrix} |\mu|^{2} + m_{H_{d}}^{2} & b \\ b & |\mu|^{2} + m_{H_{u}}^{2} \end{pmatrix} = \begin{pmatrix} m_{1}^{2} & b \\ b & m_{2}^{2} \end{pmatrix}$$

$$\sin(2\beta) = \frac{2m_3^2}{m_1^2 + m_2^2}$$

Need this to be 1!

• Of course, high-scale SUSY has been considered before

Arkani-Hamed, Dimopoulos '04 Giudice, Romanino '04

• Also, relations  $aneta \leftrightarrow \lambda(m_s) \leftrightarrow m_h$  have been discussed

cf. the 140-GeV-Higgs-mass-prediction of Hall/Nomura, '09

- <u>Our goal:</u> Identify as special structure/symmetry leading to  $\tan \beta = 1$ (i.e. to  $\lambda = 0$ )
- Indeed, such a structure is known in heterotic orbifolds:

Shift symmetry:  $K_H \sim |H_u + \overline{H}_d|^2$ 

Lopes-Cardoso, Lüst, Mohaupt '94 Antoniadis, Gava, Narain, Taylor '94 Brignole, Ibanez, Munoz, Scheich, '95...'97

In more detail:  $K_H = f(S, \overline{S})|H_u + \overline{H}_d|^2$ 

Assuming  $F_S \neq 0$  and  $m_{3/2} \neq 0$  this gives

$$m_1^2 = m_2^2 = m_3^2 = |m_{3/2} - \overline{F}^S f_{\overline{S}}|^2 + m_{3/2}^2 - F^S \overline{F}^S (\ln f)_{S\overline{S}}$$

 This shift-symmetric Higgs-Kähler potential has also been rediscovered/reused in orbifold GUTs

> K. Choi et al. '03 AH, March-Russell, Ziegler '08 Brümmer et al. '09...'10 Lee, Raby, Ratz, Ross, ... '11

• In this language, it is easy to see the physical origin:

5d SU(6)  $\rightarrow$  SU(5)×U(1); 35 = 24+5+ $\overline{5}$ +1; Higgs= $\Sigma$  +  $iA_5$ cf. Gogoladze, Okada, Shafi '07

# Comments

- This simple understanding of the shift-symmetry lets us hope that it is more generic heterotic WLs ↔ type IIA / D6-WLs ↔ type IIB / D7-WLs or positions
- These and other origins of the Higgs-shift-symmetry and of  $\tan \beta = 1$  have recently also been explored in Ibanez, Marchesano, Regalado, Valenzuela '1206...
- Clearly, we eventually need more phenomenological implications of 'stringy high-scale SUSY' (e.g. in cosmology)

Chatzistavrakidis, Erfani, Nilles, Zavala '1206... Higaki, Hamada, Takahashi '1206... Anchordoqui, Goldberg, Huang, Lüst, Taylor, Vlcek '1208...

# Corrections? Precision?

- The superpotential (e.g. top Yukawa) breaks the shift symmetry
- The crucial point is compactification

Shift symmetry is exact (gauge symmetry!) in 10d. The shift corresponds to switching on a WL. This is not a symmetry in 4d (4d-zero modes 'feel' the WL). 4d-loops destroy the shift symmetry of Kähler potential.

• Optimistic approach to estimating the 'goodness' of our symmetry:

Symmetry-violating running between  $m_c$  and  $m_s$ 

 $\Rightarrow$  Correction  $\delta \sim \ln(m_c/m_s)$ 

#### More explicitly:

$$M_{H}^{2} = (|\mu|^{2} + m_{H}^{2}) \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} + \begin{pmatrix} \delta |\mu|^{2} + \delta m_{H_{d}}^{2} & \delta b \\ \delta b & \delta |\mu|^{2} + \delta m_{H_{u}}^{2} \end{pmatrix}$$

= symmetric + loop violation

• Leading effects: *y<sub>t</sub>* and gauge

$$\delta M_H^2 = f(\epsilon_y, \epsilon_g, m_{\text{soft}})$$
 ;  $\epsilon_y = \int_{\ln m_s}^{\ln m_c} dt \, \frac{6|y_t|^2}{16\pi^2}$ 

• Enforce det  $M_H^2 = 0$  after corrections  $\Rightarrow \epsilon_y, \epsilon_g, m_{\text{soft}}$  are related

 $\cos 2\beta = \epsilon_y \times \{ \text{calculable } \mathcal{O}(1) \text{ factor} \}$ 



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# A different application of the same shift symmetry

AH, Kraus, Lüst, Steinfurt, Weigand, 1104.5016 ..., Küntzler, 1207.2766 ..., Arends, Heimpel, Mayrhofer, Schick, 12...

• Fluxbrane inflation with flat direction protected by shift symmetry for D7-brane motion



Related to WLs by mirror symmetry / T-duality

### Fluxbrane inflation

• Crucial fact: At large volume (i.e. weak flux *F*), the potential is much more flat than in brane-antibrane inflation:

$$V \sim 1 - rac{g_s}{r^{d_\perp} - 2} \quad o \quad V \sim F^2 - F^4 rac{g_s}{r^{d_\perp} - 2}$$
Hence:  $\eta \sim F^2 \ll 1$ 

• Note: This is conceptually similar to D3/D7 inflation

Dasgupta, Herdeiro, Hirano, Kallosh, '02

and T-dual to inflation from branes at angles and Wilson lines Garcia-Bellido, Rabadan, Zamora, '01 Avgoustidis, Cremades, Quevedo, '06

## Flat direction / shift symmetry

- Chose brane/bulk fluxes such that  $W_0$  does not depend on  $\varphi$ .
- Of course, since W<sub>0</sub> ≠ 0, the usual 'η-problem of supergravity' is still present:

$$\mathcal{K} = -\ln(\mathcal{S} + \overline{\mathcal{S}} + \kappa(arphi, \overline{arphi})) + \cdots \implies \eta \simeq 1 ext{ from } V_{\mathcal{F}}$$

[Here  $\kappa$  is the Kähler potential on the D7-brane moduli space; similar to situation in KKLMMT.]

- Fact: F-theory on K3×K3 has  $\kappa = \kappa(\varphi + \overline{\varphi})$
- We expect this shift-symmetric structure to arise more generally in the large complex structure limit.

Grimm, Ha, Klemm, Klevers, ....'09-'11 Alim, Hecht, Jockers, Mayr, Mertens, ....

# Conclusions / Summary

- In the absence of new electroweak physics at a TeV, the 'vacuum stability scale' ( $\lambda(\mu) = 0$ ) may be a crucial hint at new physics
- Well-motivated guess: SUSY broken with  $\tan \beta = 1$  at this scale
- Possible structural reason: shift symmetry in Higgs sector (Predictivity, i.e. m<sub>h</sub> + m<sub>t</sub> + α<sub>s</sub> ⇒ m<sub>s</sub> remains strong, even if shift symmetry is only approximate)
- The very same stringy symmetry (but in a different sector) may be crucial to maintain flatness in Fluxbrane inflation