

# Brane Moduli in Higgs Physics and Cosmology

Arthur Hebecker (Heidelberg)

- 125-GeV Higgs from a Stringy Shift Symmetry

**AH, Knochel, Weigand** (1204.2551)

- Fluxbrane Inflation with Shift symmetry

**AH, Kraus, Lüst, Steinfurt, Weigand** (1104.5016)

..., **Küntzler** (1207.2766)

..., **Arends, Heimpel, Mayrhofer, Schick** (12xx.xxxx)

## Outline

- The 125-GeV-Higgs (without SUSY) from a String-Pheno Perspective
- Main idea:  $\lambda = 0$  at some high scale (SUSY-breaking scale) due to **shift symmetry in the Higgs sector**
- Stringy (Wilson-line) origin of this shift symmetry
- Originally heterotic, but may be as natural in **D-branes**
- Closely related: The very same symmetry may be responsible for a flat potential in **fluxbrane inflation**

# Motivation

- We have a Higgs at 125 GeV and nothing else (yet?)

Of course: **low-scale SUSY is still OK**

Also: **Muon- $(g - 2)$ ;  $h \rightarrow \gamma\gamma$  excess; 130-GeV  $\gamma$ -ray line...**

- Nevertheless: What if we just had to accept the fine-tuned non-SUSY SM for a large energy range?
- Well-known: for low  $m_h$ ,  $\lambda$  runs to zero at some scale  $< M_P$  (vacuum stability bound)

Lindner, Sher, Zaglauer '89

Gogoladze, Okada, Shafi '07

...

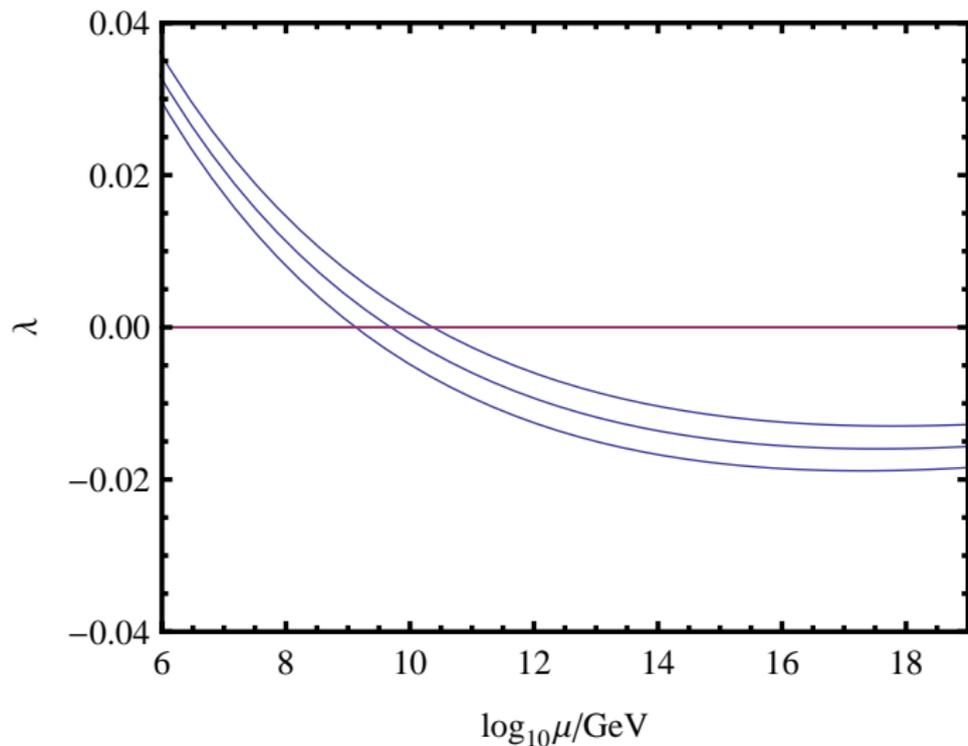
Shaposhnikov, Wetterich 09'

Giudice, Isidori, Strumia, Riotto, ...

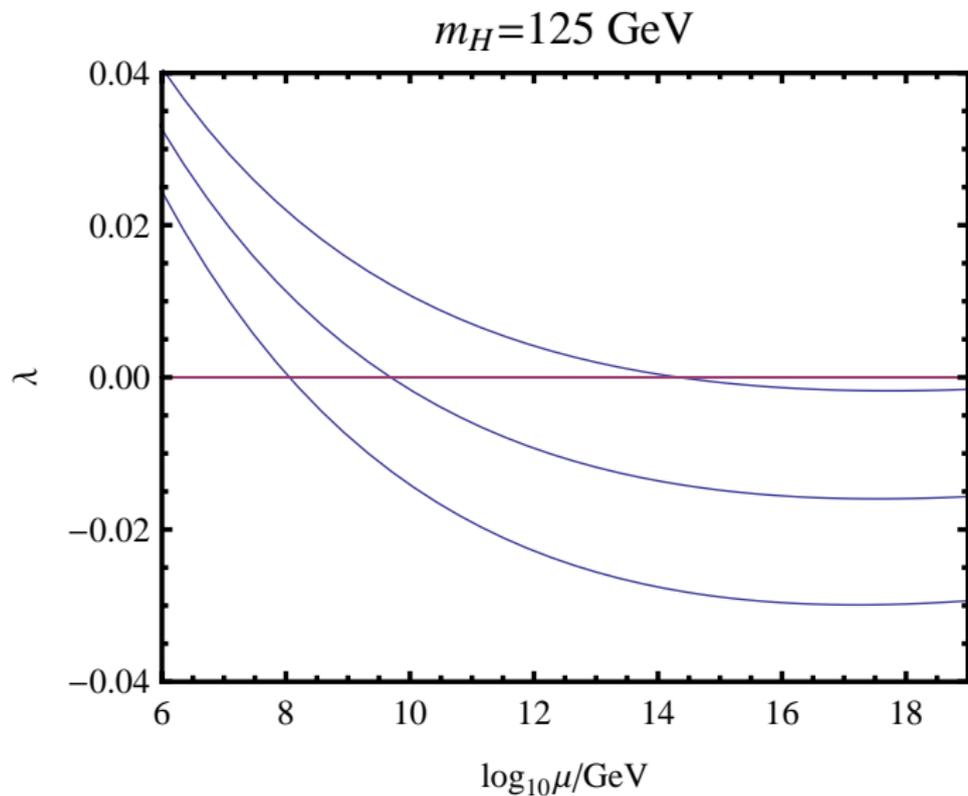
- It has been attempted to turn this into an  $m_h$  prediction

Running of  $\lambda$  (for a  $\pm 1$  GeV variation of  $m_{\text{Higgs}}$ )

$m_t = 172.9$  GeV



Running of  $\lambda$  (for a  $2\text{-}\sigma$  variation of  $m_{\text{top}}$ )



## String-phenomenologist's perspective

- Insist on stringy UV completion (for conceptual reasons)
- Expect SUSY at string/compactification scale (stability!)
- **Natural guess:** The special scale  $\mu(\lambda = 0)$  is the SUSY-breaking scale

- Crucial formula:

$$\lambda(m_s) = \frac{g^2(m_s) + g'^2(m_s)}{8} \cos^2(2\beta)$$

- Reminder:

$$M_H^2 = \begin{pmatrix} |\mu|^2 + m_{H_d}^2 & b \\ b & |\mu|^2 + m_{H_u}^2 \end{pmatrix} = \begin{pmatrix} m_1^2 & b \\ b & m_2^2 \end{pmatrix}$$

$$\sin(2\beta) = \frac{2b}{m_1^2 + m_2^2}$$

Need this to be 1!

- Of course, high-scale SUSY has been considered before

Arkani-Hamed, Dimopoulos '04  
Giudice, Romanino '04, ...

- Also, relations  $\tan \beta \leftrightarrow \lambda(m_s) \leftrightarrow m_h$  have been discussed

cf. the 140-GeV-Higgs-mass-prediction of Hall/Nomura, '09

- Our goal:

Identify as special structure/symmetry leading to  $\tan \beta = 1$   
(i.e. to  $\lambda = 0$ )

- Indeed, such a structure is known in heterotic orbifolds:

Shift symmetry:

$$K_H \sim |H_u + \bar{H}_d|^2$$

Lopes-Cardoso, Lüst, Mohaupt '94  
Antoniadis, Gava, Narain, Taylor '94  
Brignole, Ibanez, Munoz, Scheich, '95... '97  
...  
Pena, Nilles, Oehlmann, '12

**In more detail:**  $K_H = f(S, \bar{S}) |H_u + \bar{H}_d|^2$

Assuming  $F_S \neq 0$  and  $m_{3/2} \neq 0$  this gives

$$m_1^2 = m_2^2 = b = \left| m_{3/2} - \bar{F}^S f_{\bar{S}} \right|^2 + m_{3/2}^2 - F^S \bar{F}^S (\ln f)_{S\bar{S}}$$

- This shift-symmetric Higgs-Kähler potential has also been rediscovered/reused in orbifold GUTs

K. Choi et al. '03

AH, March-Russell, Ziegler '08

Brümmer et al. '09... '10

Lee, Raby, Ratz, Ross, ... '11

- In this language, it is easy to see the physical origin:

5d  $SU(6) \rightarrow SU(5) \times U(1)$ ;  $35 = 24 + 5 + \bar{5} + 1$ ; Higgs =  $\Sigma + iA_5$

cf. Gogoladze, Okada, Shafi '07

## D-brane origin

- This simple understanding of the shift-symmetry lets us hope that it is more generic

heterotic WLs  $\leftrightarrow$  type IIA / D6-WLs  $\leftrightarrow$  type IIB / D7-WLs  
or positions

- These and other origins of the Higgs-shift-symmetry and of  $\tan \beta = 1$  have recently also been explored in  
Ibanez, Marchesano, Regalado, Valenzuela '1206...

- Clearly, we eventually need **more** phenomenological implications of '**stringy high-scale SUSY**' (e.g. in cosmology)

Chatzistavrakidis, Erfani, Nilles, Zavala '1206...

Higaki, Hamada, Takahashi '1206...

Anchordoqui, Goldberg, Huang, Lüst, Taylor, Vlcek '1208...

## Corrections? Precision?

- The superpotential (e.g. top Yukawa) breaks the shift symmetry
- The crucial point is compactification

Shift symmetry is exact (gauge symmetry!) in 10d.

The shift corresponds to switching on a WL.

This is not a symmetry in 4d (4d-zero modes 'feel' the WL).

4d-loops destroy the shift symmetry of Kähler potential.

- Optimistic approach to estimating the 'goodness' of our symmetry:

Symmetry-violating running between  $m_c$  and  $m_s$

⇒ Correction  $\delta \sim \ln(m_c/m_s)$

## More explicitly:

$$M_H^2 = (|\mu|^2 + m_H^2) \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} + \begin{pmatrix} \delta|\mu|^2 + \delta m_{H_d}^2 & \delta b \\ \delta b & \delta|\mu|^2 + \delta m_{H_u}^2 \end{pmatrix}$$

=            symmetric            +            loop violation

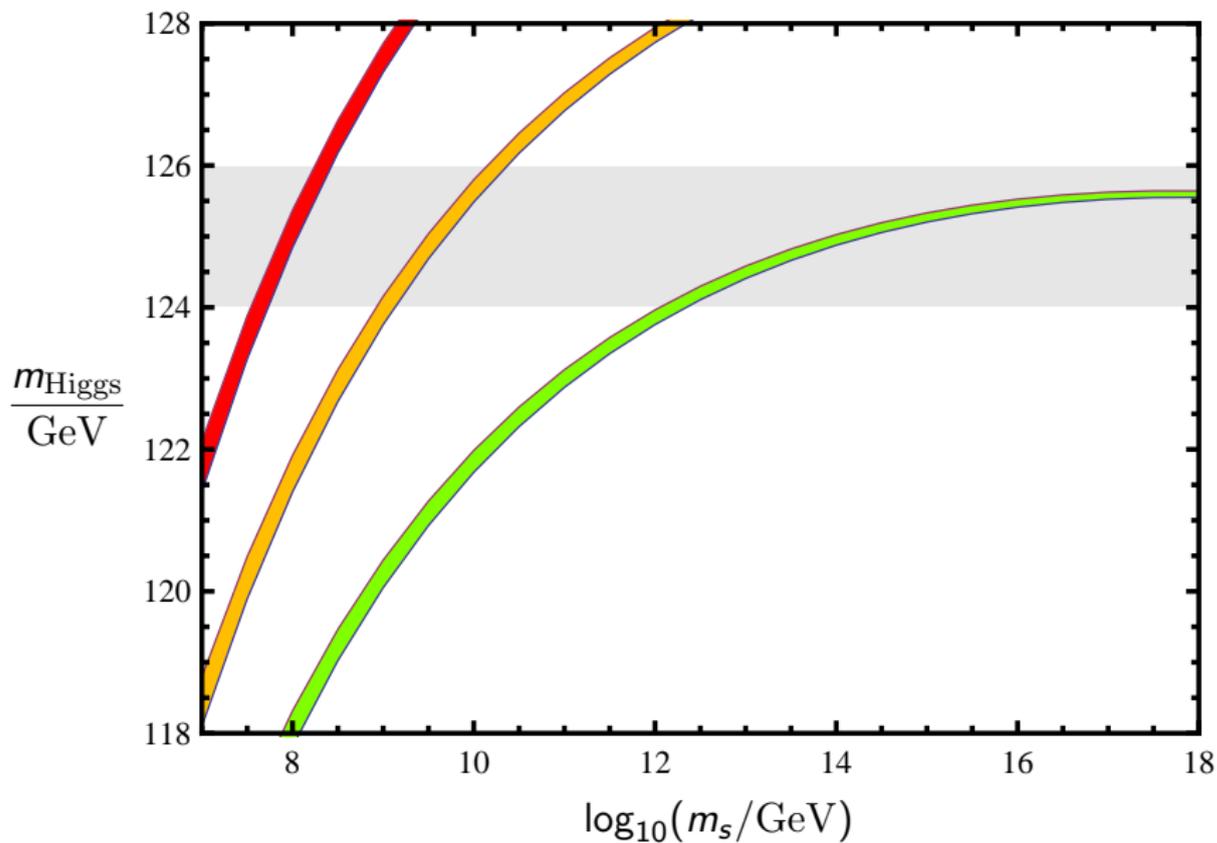
- Leading effects:  $y_t$  and gauge

$$\delta M_H^2 = f(\epsilon_y, \epsilon_g, m_{\text{soft}}^i) \quad ; \quad \epsilon_y = \int_{\ln m_s}^{\ln m_c} dt \frac{6|y_t|^2}{16\pi^2}$$

- Enforce  $\det M_H^2 = 0$  after corrections  $\Rightarrow \epsilon_y, \epsilon_g, m_{\text{soft}}^i$  are related

$$\cos 2\beta = \epsilon_y \times \{\text{calculable } \mathcal{O}(1) \text{ factor}\}$$

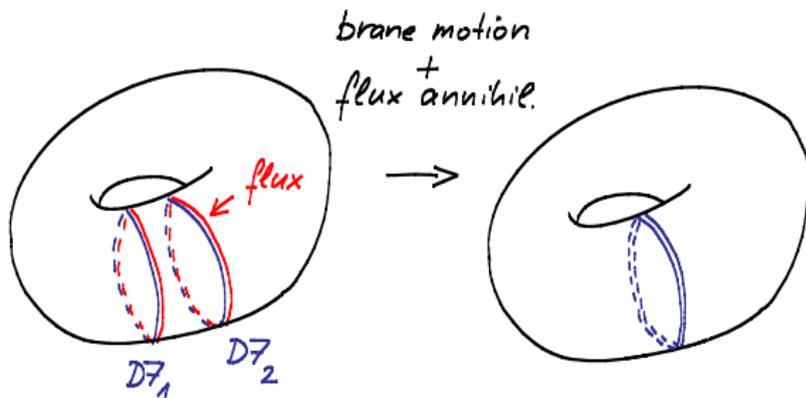
Assumption:  $(m_s < m_c < 100m_s)$



# A different application of the same shift symmetry

AH, Kraus, Lüst, Steinfurt, Weigand, 1104.5016  
..., Küntzler, 1207.2766  
..., Arends, Heimpel, Mayrhofer, Schick, 12...

- Fluxbrane inflation with flat direction protected by shift symmetry for D7-brane motion



- Related to WLs by mirror symmetry / T-duality

## Fluxbrane inflation

- **Crucial fact:** At large volume (i.e. weak flux  $F$ ), the potential is much more flat than in brane-antibrane inflation:

$$V \sim 1 - \frac{g_s}{r^{d_\perp - 2}} \quad \rightarrow \quad V \sim F^2 - F^4 \frac{g_s}{r^{d_\perp - 2}}$$

Hence:  $\eta \sim F^2 \ll 1$

- **Note:** This is conceptually similar to [D3/D7 inflation](#)

Dasgupta, Herdeiro, Hirano, Kallosh, '02

and T-dual to inflation from [branes at angles](#) and [Wilson lines](#)

Garcia-Bellido, Rabadan, Zamora, '01  
Avgoustidis, Cremades, Quevedo, '06

## Flat direction / shift symmetry

- Chose brane/bulk fluxes such that  $W_0$  does not depend on  $\varphi$ .
- Of course, since  $W_0 \neq 0$ , the usual 'η-problem of supergravity' is still present:

$$K = -\ln(S + \bar{S} + \kappa(\varphi, \bar{\varphi})) + \dots \quad \implies \quad \eta \simeq 1 \text{ from } V_F$$

[Here  $\kappa$  is the Kähler potential on the D7-brane moduli space; similar to situation in KKLMNT.]

- **Fact:** F-theory on  $K3 \times K3$  has  $\kappa = \kappa(\varphi + \bar{\varphi})$
- We expect this **shift-symmetric** structure to arise more generally in the **large complex structure limit**.

Grimm, Ha, Klemm, Klevers, ... '09-'11  
Alim, Hecht, Jockers, Mayr, Mertens, ...

## Conclusions / Summary

- In the absence of new electroweak physics at a TeV, the 'vacuum stability scale' ( $\lambda(\mu) = 0$ ) may be a crucial hint at new physics
- Well-motivated guess: SUSY broken with  $\tan \beta = 1$  at this scale
- Possible structural reason: shift symmetry in Higgs sector
- Interesting task: Explicit realization in D-brane models
- The very same stringy symmetry (but in a different sector) may be crucial to maintain flatness in Fluxbrane inflation